

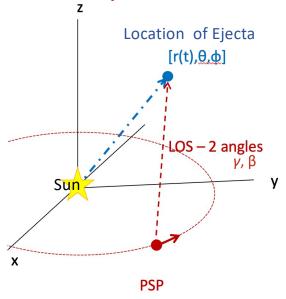
STEREO-A/SECCHI and PSP/WISPR Joint Observations of Solar Minimum CMEs



P. Liewer, C. Braga, J. R. Hall, P. Hess, R. A. Howard, P. Penteado, J. Qiu, G. Stenborg, STEREO Team Meeting, 18 January 2022

- Tracking & Fitting Technique for determining the trajectories of solar eject from WISPR images that takes into account PSP's rapid motion
- CME from Orbit 2 (2019 Apr 2) STEREO HI-1A data used to verify Tracking & Fitting Technique (*Liewer et al Sol Phys 2020*)
- Streamer Blowout CME from Orbit 4, 25-27 Jan 2020 (Liewer et al A&A 2021)
 - Used simultaneous ST/ HI-1A & PSP observation to triangulate to determine CME location
 - Unexpected evolution of magnetic flux rope captured by STA/EUVI Is it the source?
- CME from Orbit 7 (20 Jan 2021)
 - Evolution of high latitude prominence eruption captured by STA EUVI, COR1, COR2
 - WISPR images suggest a nested flux rope structure of CME

Technique for Tracking & Fitting Ejecta for Trajectory Determination



- Ejecta's pixel location in an image defines a line-of-sight from PSP, described by two angles referenced to WISPR point-of-view
 - We use angles defined relative to PSP orbit plane in & out of the plane $\gamma(t)$, $\beta(t)$
- Assume ejecta moves radially at constant velocity => ejecta's longitude and latitude are constant in the Heliocentric Inertia (HCI) frame
- Derived 2 equations to relate ejecta's HCI coordinates $[r(t), \theta, \phi]$ to image coordinate angles $\gamma(t)$, $\beta(t)$
- Developed a procedure for fitting measured $\gamma(t)$, $\beta(t)$, obtained from a sequence of images, to the 2 equations
- Manually track a feature in a sequence of images to obtain a set $[\gamma(t_i), \beta(t_i)]$
- Fit the data to the 2 equations to determine the 4 unknown trajectory parameters: longitude, latitude, velocity, and radius (distance from the Sun) at the start of tracking.
- Tracking & Fitting method described in Liewer et al, Sol Phys 2020

3D Trajectory of 2019 April 2 "Skull" CME (PSP \cong 40 R_{sun})

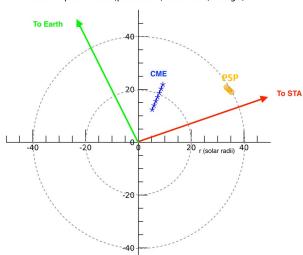
• From Tracking (Lower Eye) and Fitting, we determine the 3D trajectory of a CME feature (Lower Eye) in HCI coordinates (*Liewer et al, Sol Phys 2020*):

HCI longitude = 66°; Latitude = 6°, V = 333 km/s; $R(0)/R_{sun} = 13.4 @ t_0 = 2019 Apr 2 12:09 UTC$

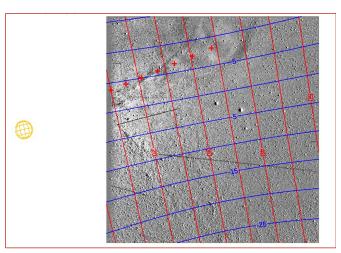
WISPR-I (r-difference)

Polar plot of 3D trajectory points HCl coordinate system

2019 April 2 CME (predicted) and PSP (orange)



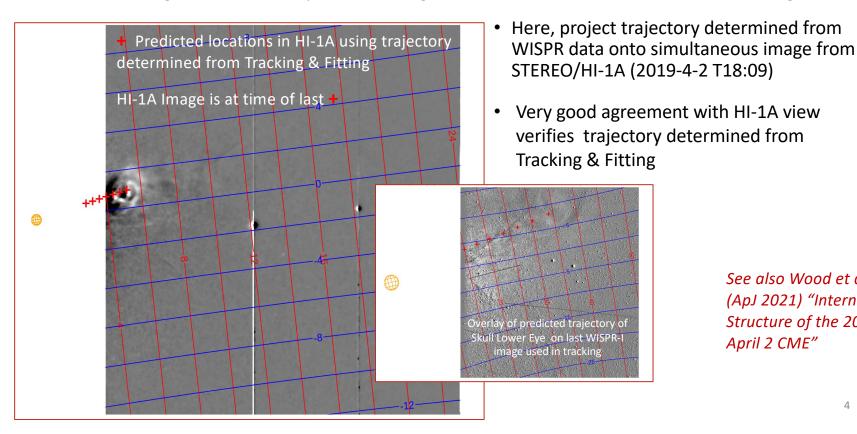
Overlay of predicted trajectory of Skull Lower Eye on last WISPR-I image used in tracking



• Can we use STEREO-A view of CME to verify trajectory?

Verify Tracking & Fitting Result using STEREO A Viewpoint

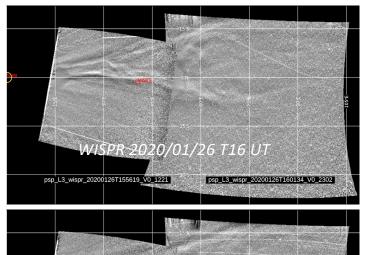
• Developed projection software to take 3D trajectory points HCI (x, y, z) and re-projects them to proper location in image from another spacecraft using WCS information in FITS headers of the image



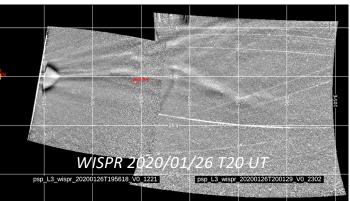
See also Wood et al (ApJ 2021) "Internal Structure of the 2019 April 2 CME"

Orbit 4: CME seen by STEREO A on 25 Jan 2020

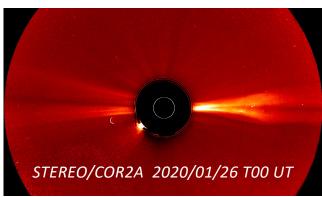
Identified as Streamer Blowout (SBO) CME - Cavity, with no bright leading edge, pushes streamer apart --- CME followed by streamer evacuation



- Both see only cavity
 no bright leading
 edge
- Both see streamer pushed apart by CME cavity



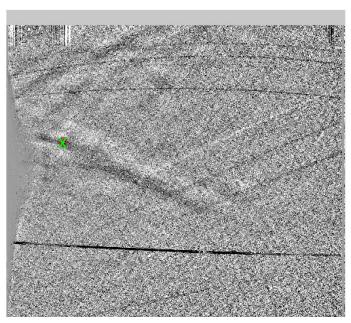
- CME shows similar V-shaped structure in HI-1A
- HI-1 shows more outflow behind CME
 - STA/PSP Separation angle ~50° 60°





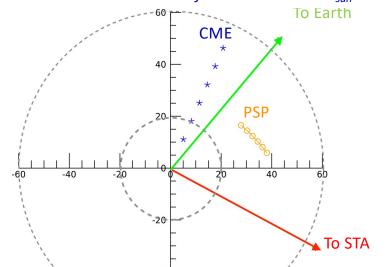
26-27 Jan 2020 CME Trajectory from Tracking & Fitting

- Here, tracked a dark spot at behind CME cavity in WISPR-O for 10 hrs
- T&F solution for trajectory in HCl coordinates: $(r, \theta, \phi) = (30 R_{sun}, 65^{\circ}, 2^{\circ})$
 - Recall, Tracking & Fitting assumes radial motion and a constant velocity



WISPR-O rdiff image 2020 Jan 26 @ 20:49 PSP ~30 R_{sun} from the Sun





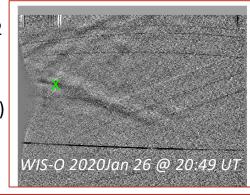
- STEREO A also saw the CME 50 ° separation
- Can we verify trajectory using STEREO A?

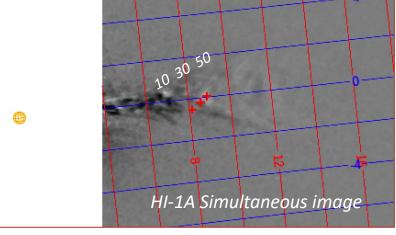
Verify CME Trajectory by Triangulation between STA & WISPR

Used projection software to take 3D locations HCI (x, y, z) and re-project onto image from STEREO HI-1A using WCS information in FITS header

- Pixel X on CME 's Feature in WISPR image defines a LOS through space (2 angles)
- Feature can be anywhere along the LOS
- Choosing a distance R_{psp} from SC along LOS gives a 3D HCI location (x, y, z)
- Project onto simultaneous FITS image from HI1-A and iterate to find distance R_{psp} such that projected point falls on same feature
- Right Image: Location of WISPR feature X projected onto simultaneous HI-1A image for 3 distances R_{psp} from PSP: 10, 30 & 50 R_{sun}
- Point for R_{psp}= 30 falls near same feature in HI-1
- Iterating, found HCI coordinates (R, lon, lat)= (31±2, 66°±3,-2°±2, in excellent agreement with trajectory found from Tracking & Fitting

Triangulation makes NO assumption about the motion, but limited by ability to identify same feature in both images





What was the source of 2020 Jan SBO CME?

STEREO-A Observations of Flux Rope Evolution

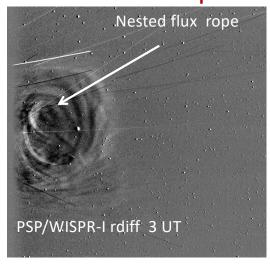
- EUVI 171 shows evidence of Flux Rope formation & rise
 23 Jan 12-24 UT
- STA Data gap 24 Jan (pause in movie)
- Structure in same location leaves EUVI FOV 25 Jan ~6 UT.
- SBO CME seen in COR2A ~13 UT
- Exhaustive search of LASCO, WISPR shows no sign of ejection on 24 Jan
- Search of HI-1&2 for 25 Jan showed no suggestion of CME having left Sun on 24 Jan

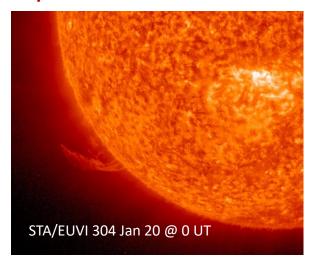
Watch here! EUVI A (1711): 20200123 095800 UT

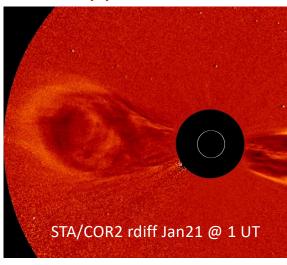
Conclude:

- CME source is flux rope constrained in corona from formation on Jan 23 to ejection on Jan 25
- Trajectory from Tracking & Fitting (verified with by Triangulation with HI-1A) give longitude inconsistent with alternative AR source HCI longitude is 40° west of AR location on Jan 25 (*Liewer et al A&A 2021*)

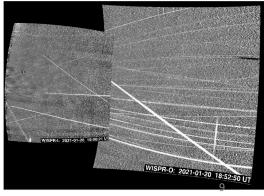
Orbit 7: 2021 January 20 Filament Eruption and CME seen by STEREO A PSP & STA Separated by 185°-- WISPR sees CME from opposite side



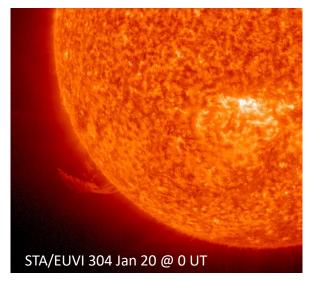


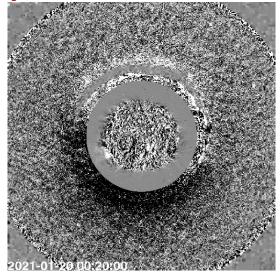


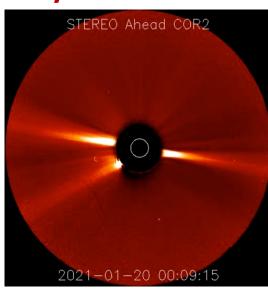
- Trajectory determination shows CME roughly midway between PSP and STA
- Both WISPR and COR2 show flux rope structure WISPR shows more internal structure suggesting nested flux ropes (Veronig et al 2018)
- Is the "eye" the original flux rope seen bey STA/EUVI?
- Note: After leaves COR2A FOV, dramatic changes in shape of CME seen by WISPR (Braga et al, in preparation)



Orbit 7 2021 January 20 CME Evolution Seen by STEREO A



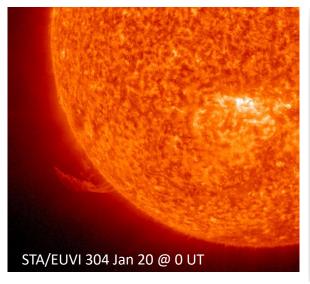


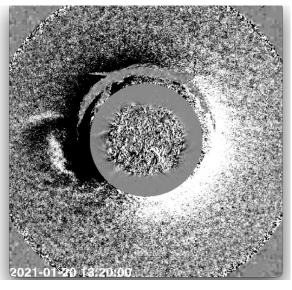


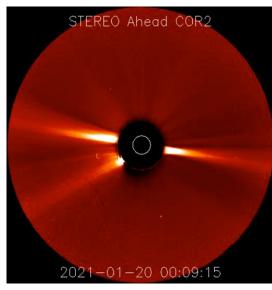
Composite by P. Hess

- Late 1/20 to early 1/21: EUVI 304 observes rise of prominence ~60 S just behind east limb
- 1/21 6-12 UT EUVI-COR1 movie shows cavity expanding asymmetrically to the north
 - Presumbably, CME is guided to the HCS by the overlying magnetic fields (*Byrne, 2010; Liewer et al, 2015*)
- Jan 21 12UT Tip of cavity visible in COR2 slightly north of equator

Orbit 7 2021 January 20 CME Evolution Seen by STEREO A







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- Joint observations of 2010 Apr 2 CME: STEREO HI-1A data used to verify trajectory determination using WISPR data alone
 - New Tracking & Fitting Technique for determining the trajectories of solar eject needed because
 of effect of PSP's rapid motion on image analysis (Liewer et al Sol Phys 2020)
- Streamer Blowout CME 25-27Jan2020 (Liewer et al A&A 2021)
 - Used simultaneous STA/HI-1 & PSP observation to triangulate to determine CME location
 - Results Consistent with trajectory solution from Tracking & Fitting of WISPR data
 - From STA/EUVI 171, concluded that CME source is probably flux rope constrained in corona from formation on Jan 23 to ejection on Jan 25
- CME from Orbit 7 January 25-27, 2021
 - Evolution of from high latitude prominence to low latitude CME seen by SECCHI EUVI, COR1&2
 - Cavity seen to expand asymmetrically as CME is guided to HCS by overlying magnetic field
 - WISPR images suggest a nested flux rope structure of CME